Weak solutions to the isentropic system of gas dynamics

Cheng Yu
in collaboration with
Ming Chen and Alexis Vasseur

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Outline

- 1. Introduction
 - ► The governing equations
 - weak solutions to the Incompressible flows

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 - ► The governing equations
 - weak solutions to the Incompressible flows
- 2. Construction of ∞ many global admissible weak solutions
 - Main result
 - Key idea and steps

The governing equations

- Conservation Laws.
- ► Navier-Stokes: 1827-1845.

$$\begin{split} \rho_t + \operatorname{div}(\rho \mathbf{u}) &= 0, \\ (\rho \mathbf{u})_t + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u}) &= \underbrace{-\nabla P}_{\text{preasure}} + \underbrace{\operatorname{div}(\mu \mathbb{D}(\mathbf{u})) + \nabla(\lambda \operatorname{div} \mathbf{u})}_{\text{diffusion.}} \end{split}$$

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- $\mathbb{D}(\mathbf{u}) = \frac{1}{2}(\nabla \mathbf{u} + \nabla^T \mathbf{u}), P = \alpha \rho^{\gamma}.$
- \triangleright μ is the bulk viscosity and λ is the shear viscosity.

The governing equations

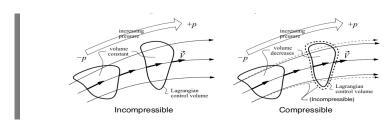
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- \blacktriangleright μ is the bulk viscosity and λ is the shear viscosity.
- ► If $\mu = \lambda = 0$, the Euler equations (1757).
- For some fluids and gases, experiments indicate that the viscous influence is very small. This is why it is often set to $\mu = \lambda = 0$ for some fluids, which is called the inviscid case.

Incompressibility condition

► Incompressible flow: The volume does not change in time.



From Mark Drela's lecture note on Fluid Mechanics. $\mathbf{u} = 0$ iff $\det J = 1$ iff incompressible flow.

▶ If ρ = constant, then

$$\mathbf{u}_t + \mathbf{u} \cdot \nabla \mathbf{u} + \nabla P - \mu \Delta \mathbf{u} = 0,$$

$$\operatorname{div} \mathbf{u} = 0.$$

Question??

- ► A longstanding open question has been to determine what degree(s) of regularity must assumed to guarantee:
 - uniqueness (still opening)
 - conservation of energy for weak solutions (Onsager's conjecture, proved).
 - Uniqueness or conservation of energy fails when the degree below a critical value.
- Question: Can we construct +∞ many global admissible weak solutions?
 - Under admissible condition??
 - It will narrow down further the class of weak solutions to single out physical relevant solutions of the Euler equations for the uniqueness.

Incompressible Euler equations

▶ A solution (**u**,*P*) to the incompressible Euler equations is such that

$$\begin{cases} \partial_t \mathbf{u} + \operatorname{div} (\mathbf{u} \otimes \mathbf{u}) + \nabla P = 0, & x \in \mathbb{T}^3, \\ \operatorname{div} \mathbf{u} = 0. \end{cases}$$

If the solution is sufficiently smooth, say \mathbb{C}^1 , then the total kinetic energy

$$E(\mathbf{u}) := \frac{1}{2} \int_{\mathbb{T}^3} |\mathbf{u}(t, x)|^2 dx$$

is conserved, and any solution is uniquely determined by the initial data.

► A folklore conjecture: Uniqueness should fail when $\mathbf{u} \in C^{\alpha}$ for some $\alpha < 1$, which is highly linked to Onsager's conjecture.

Onsager's semi-formal proof of the sufficient condition

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Onsager's semi-formal proof of the sufficient condition

- Roughly speaking, enough regularity allows us to control convective term and to do integration by parts.
- ► The term to control is the total energy flux

$$\Pi = \langle \operatorname{div}(\mathbf{u} \otimes \mathbf{u}), \mathbf{u} \rangle \sim \left\langle (\nabla^{1/3} \mathbf{u} \otimes \nabla^{1/3} \mathbf{u}) : \nabla^{1/3} \mathbf{u} \right\rangle$$

Thus the quantity $\|\nabla^{1/3}\mathbf{u}\|_{L^3}$ appears. Any better regularity would be sufficient to justify integration by parts to show that the flux $\Pi = 0$.

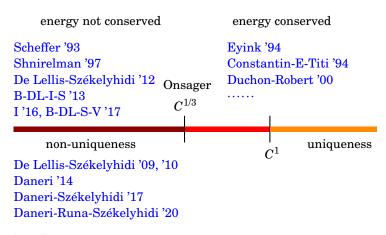
Onsager's Conjecture [Onsager '49]

The threshold Hölder regularity for the validity of the energy conservation of weak solutions has exponent 1/3:

- (1) Every weak solution **u** to the Euler equations with Hölder continuity exponent $\alpha > \frac{1}{3}$ conserves energy.
- (2) For any $\alpha < \frac{1}{3}$ there exists a weak solution $\mathbf{u} \in C^{\alpha}$ which does not conserve energy.



Threshold regularity



based on a Baire category argument

Weak solutions to the Cauchy problem

$$\left\{ \begin{array}{l} \partial_t \mathbf{u} + \mathrm{div} \; (\mathbf{u} \otimes \mathbf{u}) + \nabla p = 0, \quad x \in \Omega, \\ \mathrm{div} \mathbf{u} = 0, \\ \mathbf{u}|_{t=0} = \mathbf{u}^0. \end{array} \right.$$

A divergence free vector field $\mathbf{u} \in L^{\infty}_t L^2_x$ is a *global admissible* weak solution if

- $\int_0^\infty \int_\Omega \left(\mathbf{u} \cdot \partial_t \varphi + \mathbf{u} \otimes \mathbf{u} : \nabla \varphi \right) dx dt = -\int_\Omega \mathbf{u}^0 \cdot \varphi(\cdot, 0) dx$ for every test function $\varphi \in C_c^\infty$ with $\operatorname{div} \varphi = 0$.
- $\int_{\Omega} \frac{1}{2} |\mathbf{u}(\cdot,t)|^2 dx \le \int_{\Omega} \frac{1}{2} |\mathbf{u}^0(\cdot)|^2 dx \quad \text{for every } t \ge 0.$

Non-uniqueness and density of 'wild' data

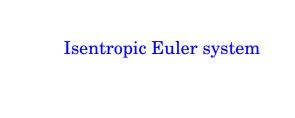
Theorem (Székelyhidi-Wiedemann '12, Chen-Vasseur-Y.)

For any $\varepsilon > 0$ and any $\mathbf{u}^0 \in L^2(\mathbb{T}^n)$, there exist infinitely many $v^0 \in L^2(\mathbb{T}^n)$ satisfying

$$\|v^0 - \mathbf{u}^0\|_{L^2(\mathbb{T}^n)}^2 < \varepsilon,$$

such that for each such initial value v^0 , there exist infinitely many global admissible weak solutions v to the incompressible Euler equations.

- Construct a sub-solution by vanishing viscosity limit from Navier-Stokes.
 - Leray-Hopf theory for N.-S.
 - Euler equations: No results of global existence of weak solutions.
 - Inviscid limit ($\mu \rightarrow 0$): weak limit is not commutative with nonlinear term.
- Partial Parti



Weak solutions

$$\begin{split} & \rho_t + \operatorname{div}(\rho \mathbf{u}) = 0, \\ & (\rho \mathbf{u})_t + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u}) + \nabla P = 0 \end{split}$$

$$\begin{split} \int_0^\infty \int_\Omega \left(V \cdot \partial_t \varphi + \frac{V \otimes V}{\rho} : \nabla \varphi + \rho^\gamma \operatorname{div} \varphi \right) dx dt \\ &= - \int_\Omega V^0 \cdot \varphi(\cdot, 0) dx \end{split}$$

where $V = \rho \mathbf{u}$.

Related works

- The proof relies on the *Convex integration* machinery developed by De Lellis–Székelyhidi.
- Two directions of the isentropic flow
 - One direction, pioneered by Chiodaroli, considers a wide class of initial densities. Some extensions, Luo-Xie-Xin, and Feireisl.
 - ► The other direction, pioneered by Chiodaroli–De Lellis–Kreml, focuses on initial values being Riemann data.
 - Extensions of both strategies have been studied for the full Euler system, see Chiodaroli-Feireisl-Kreml, Al
 Baba-Klingenberg-Kreml-Mácha-Markfelder.
- Without energy condition, non-unique solutions can be constructed for any fixed initial values, see Abbatiello-Feireisl.
- A natural problem consists in studying the size of the class of initial values leading to non-unique solutions.

Riemann data

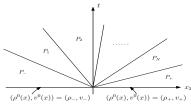
Theorem (Chiodaroli- De Lellis-Kreml, CPAM.)

For $\gamma = 2$ in 2D, there are infinitely many bounded admissible solutions with the initial data

$$(\rho^0, \boldsymbol{u}^0) = \begin{cases} (\rho_-, \boldsymbol{u}_-), & if \ x_2 < 0 \\ (\rho_+, \boldsymbol{u}_+), & if \ x_2 > 0. \end{cases}$$

- Admissible condition: energy inequality in distribution sense.
- Initial data is Riemann data.
- Key idea: sub-solutions+ convex integral.

Key idea of CDK



- ► Classical theory in 1D conservation laws: Rankine-Hugoniot conditions.
- Sub-solutions: $(\bar{\rho}, \bar{\mathbf{u}}) = \sum_{-}^{+} (\rho, \mathbf{u}) \mathbb{I}_{P_i}$
- Oscillation lemma: Let $\bar{\mathbf{u}} \otimes \bar{\mathbf{u}} \bar{U} < \frac{C}{n} Id$, there exists infinitely many bounded maps $(\underline{\mathbf{u}}, \underline{U}) \in L^{\infty}$, such that
 - $\underline{\mathbf{u}}$, \underline{U} vanish identically outside Ω ,

 - $(\bar{\mathbf{u}} + \underline{\mathbf{u}}) \otimes (\bar{\mathbf{u}} + \underline{\mathbf{u}}) (\bar{U} + \underline{U}) = \frac{c}{n} Id.$
- Solutions: $(\rho, \mathbf{u}) = (\bar{\rho}, \bar{\mathbf{u}} + \underline{\mathbf{u}}).$

Our further understanding from CDK

- Note that $\mathbf{u} = \bar{\mathbf{u}} + \mathbf{u} = \text{mean flow} + \text{fluctuation}$.
- ► This motivates us to reformulate the system for sub-solutions as

$$\begin{split} & \rho_t + \mathrm{div}(\rho \bar{\mathbf{u}}) = 0, \\ & (\rho \bar{\mathbf{u}})_t + \mathrm{div}(\rho \bar{\mathbf{u}} \otimes \bar{\mathbf{u}} + \bar{P}I_n + \rho R) = 0. \end{split}$$

where the Reynolds stress

$$R = \overline{\mathbf{u} \otimes \mathbf{u}} - \overline{\mathbf{u}} \otimes \overline{\mathbf{u}} + (\overline{\rho^{\gamma}} - \overline{\rho}^{\gamma}) I_n$$

is symmetric and positive semidefinite.

Main result

Theorem (Chen-Vasseur-Y., Adv. Math, 2021)

Whenever $1 < \gamma \le 1 + \frac{2}{n}$, for any $\varepsilon > 0$ and any (ϱ^0, U^0) such that $E(\rho^0, U^0) \in L^1(\mathbb{T}^n)$, there exist infinitely many (ρ^0, V^0) satisfying $\rho^0 > 0$. $E(\rho^0, V^0) \in L^1(\mathbb{T}^n)$,

$$\left\| v^0 - \rho^0 \right\|^{\gamma} + \left\| \frac{V^0}{V^0} - \frac{U^0}{V^0} \right\|^2 \le \varepsilon$$

$$\|\rho^0 - \varrho^0\|_{L^{\gamma}(\mathbb{T}^n)}^{\gamma} + \left\| \frac{V^0}{\sqrt{\rho^0}} - \frac{U^0}{\sqrt{\varrho^0}} \right\|_{L^2(\mathbb{T}^n)}^2 < \varepsilon,$$

such that, for each of such initial values (ρ^0, V^0) , there exist infinitely many global admissible weak solutions (ρ, V) to the isentropic Euler equations.

Remarks: ∞ many solutions

- The most interesting range of γ in physics is $1 < \gamma \le \frac{5}{3}$ in 3D.
- ► This result can be regarded as a compressible counterpart of the one obtained by Szekelyhidi–Wiedemann (ARMA, 2012) for incompressible flows.
- ► The admissibility condition is defined in its integral form. In particular, the energy is decreasing in time *t*.
- The energy equality could be hold under particular conditions, see Y.(ARMA,2017), R. Chen-Y.(JMPA,2019), Akramov-Debiec-Skipper-Wiedemann (Anal. PDE, 2020), Feireisl-Gwiazda-Swierczewska-Gwiazda-Wiedemann(ARMA,2017)

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Key steps

- Two steps: the construction of *subsolutions*, and the convex integration of these subsolutions to obtain actual solutions.
- Can we construct a sub-solution as follows

$$\begin{split} & \rho_t + \operatorname{div}(\rho \mathbf{u}) = 0, \\ & (\rho \mathbf{u})_t + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u} + P(\rho)I_n + \rho R) = 0? \end{split}$$

- Vanishing viscosity limits from the Navier-Stokes equation.
- Weak limits for nonlinear term can produce R.
- We need a suitable convex integral tool?
 - a topological Bairé category argument.
- The energy-compatible subsolution (ρ, V, R) , denoting $U := (V \otimes V Id|V|^2/n)/\rho$, the oscillatory perturbations $(\widetilde{V}, \widetilde{U})$, readily generate $(\rho, V + \widetilde{V})$ as solutions to the the isentropic Euler system.

Existence of NS

Proposition

Fo any $\gamma > 1$, there exists the global weak solution (ρ_{ν}, V_{ν}) to

$$\label{eq:continuous_equation} \begin{cases} \partial_t \rho_{\scriptscriptstyle V} + \operatorname{div} V_{\scriptscriptstyle V} = 0, \\ \\ \partial_t V_{\scriptscriptstyle V} + \operatorname{div} \left(\frac{V_{\scriptscriptstyle V} \otimes V_{\scriptscriptstyle V}}{\rho_{\scriptscriptstyle V}} + p(\rho) I_n \right) = \operatorname{div} \left(\sqrt{v \rho_{\scriptscriptstyle V}} \mathbb{S}_{\scriptscriptstyle V} \right), \end{cases}$$

$$where \sqrt{v\rho_{\mathcal{V}}} \mathbb{S}_{\mathcal{V}} := v\rho_{\mathcal{V}} \mathbb{D} v_{\mathcal{V}} \quad with \quad \mathbb{D} v_{\mathcal{V}} := \left(\frac{\nabla v_{\mathcal{V}} + \nabla^T v_{\mathcal{V}}}{2}\right) \quad and \quad V_{\mathcal{V}} = \rho_{\mathcal{V}} v_{\mathcal{V}}.$$

- ► This weak solution was constructed by Vasseur-Y. and Bresch-Vasseur-Y. .
- The standard theory need $\gamma > \frac{3}{2}$ in the framework of Lions-Feireisl.
- The most interesting range of γ in physics is $1 \le \gamma \le \frac{5}{3}$.

Vanishing viscosity limits

As $v \rightarrow 0$, up to a subsequence,

$$(\rho_{\nu}, V_{\nu}) \rightarrow (\rho, V)$$
 weakly in $L^{\infty}(\mathbb{R}_+; L^{\gamma}(\mathbb{T}^n)) \times L^{\infty}(\mathbb{R}_+; L^{\frac{2\gamma}{\gamma+1}}(\mathbb{T}^n)),$

which defines

$$R := \lim_{v \to 0} \frac{v_v \otimes v_v}{\rho_v} - \frac{v \otimes v}{\rho}, \qquad r := \lim_{v \to 0} p(\rho_v) - p(\rho) \quad \text{in } \mathscr{D}'.$$

 $ightharpoonup \frac{|V_{\nu}|^2}{\rho_{\nu}}
ightharpoonup \frac{|V|^2}{\rho} + \operatorname{Tr}R, \ P(\rho_{\nu})
ightharpoonup P(\rho) + r,$ by energy inequality, we have

$$\int_{\mathbb{T}^n} \left(E(\rho,V) + \frac{1}{2} \mathrm{Tr} R + \frac{r}{\gamma-1} \right) dx \leq E_0.$$

Then there exist a subsolution (ρ, V, R, r) of the compressible Euler equations with energy inequality, called (\mathcal{E}^0, T) -energy compatible subsolution.

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- Smoothing via convolution.
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- The above two procedures respect energy compatibility because of convexity.
- Therefore we are left to consider convex integration from smooth energy compatible subsolutions with positive definite total defect matrix $R + rI_n$.

Oscillation lemma

Proposition (Chen-Vasseur-Y., 2021.)

There exist infinitely many \widetilde{V} and traceless \widetilde{U} (as oscillatory perturbations), both supported in Ω , such that in $\mathbb{R}^n \times \mathbb{R}_+$:

$$\begin{cases} \operatorname{div} \widetilde{V} = 0, \\ \partial_t \widetilde{V} + \operatorname{div} \widetilde{U} = 0, \end{cases}$$

while

$$\frac{(V+\widetilde{V})\otimes (V+\widetilde{V})}{\rho} - (U+\widetilde{U}) = \left(\frac{|V|^2}{n\rho} + q\right)I_n$$

is achieved as to eliminate the Reynolds stress $R := qI_n$.

Energy injection

 $(\rho, V + \tilde{V})$ Euler solution.

$$\frac{|V + \tilde{V}|^2}{\rho} = \frac{|V|^2}{\rho} + \text{tr}R.$$

 $\frac{1}{2}$ trR is pumped into the kinetic energy density through C.I..

The subsolutions

$$\begin{cases} \partial_t \rho + \operatorname{div} V = 0, \\ \partial_t V + \operatorname{div} \left(\frac{V \otimes V}{\rho} + p(\rho) Id + R \right) = 0. \end{cases}$$

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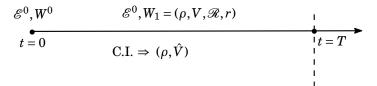
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The energy-compatible subsolution (ρ, V, R) , denoting $U := (V \otimes V - Id|V|^2/n)/\rho$, the oscillatory perturbations $(\widetilde{V}, \widetilde{U})$, readily generate $(\rho, V + \widetilde{V})$ as solutions to the the isentropic Euler system.

$$\begin{split} \mathcal{E}^0, W^0 &= (\rho^0, V^0, \mathcal{R}^0, r^0) \Rightarrow (\mathcal{E}^0, T) \text{-compatible subsolution } W_1 \\ \text{Say} \quad \rho > 0, \quad R &= \mathcal{R} + r \mathbf{I}_n > 0 \end{split}$$



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$$\mathcal{E}^{0}, W^{0} \qquad \mathcal{E}^{0}, W_{1} = (\rho, V, \mathcal{R}, r)$$

$$t = 0 \qquad \qquad t = T$$

$$\text{Energy} \quad E(\rho, \hat{V}) = E(\rho, V) + \frac{1}{2} \text{tr} R$$

$$= E(\rho, V) + \frac{1}{2} \text{tr} \mathcal{R} + \frac{n}{2} r$$

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$$\text{Energy} \neq : \quad E(\rho, \hat{V}) \text{ v.s. } E(\rho, V) + \frac{1}{2} \text{tr} \mathcal{R} + \frac{r}{r-1}$$

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$$1 < \gamma \le 1 + \frac{2}{n}$$

 \Rightarrow need compensation for

potential energy density

$$\mathcal{E}^{0}, W^{0} = (\rho^{0}, V^{0}, \mathcal{R}^{0}, r^{0}) \Rightarrow (\mathcal{E}^{0}, T) \text{-compatible subsolution } W_{1}$$

$$\text{Say} \quad \rho > 0, \quad R = \mathcal{R} + r \mathbf{I}_{n} > 0$$

$$\mathcal{E}^{0}, W^{0} \qquad \qquad \mathcal{E}^{0}, W_{1} = (\rho, V, \mathcal{R}, r)$$

$$t = 0 \qquad \qquad t = 0$$

$$r_{c}(t) = \left(\frac{2}{n(\gamma - 1)|\mathbb{T}^{n}|}\right) \int_{\mathbb{T}^{n}} r(t, x) dx$$

$$\text{Consider } (\mathcal{E}^{0}, T) \text{-compatible subsolution}$$

$$W = (\rho, V, \mathcal{R}, r + r_{c})$$

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$$\mathcal{E}^{0}, W^{0} \qquad \mathcal{E}^{0}, W_{1} = (\rho, V, \mathcal{R}, r)$$

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Consider (\mathcal{E}^{0}, T) -compatible subsolution

$$W = (\rho, V, \mathcal{R}, r + r_c)$$

Issue: bump-up of initial energy

$$E(\rho^0,V^0) \longrightarrow E(\rho^0,V^0) + \tfrac{1}{2} \mathrm{tr} R$$

$$\mathcal{E}^{0}, W^{0} = (\rho^{0}, V^{0}, \mathcal{R}^{0}, r^{0}) \Rightarrow W = (\rho, V, \mathcal{R}, r + r_{c})$$
Say $\rho > 0$, $R = \mathcal{R} + r\mathbf{I}_{n} > 0$

$$\mathcal{E}^{0}, W^{0} \qquad \qquad \mathcal{E}^{0}, W = (\rho, V, \mathcal{R}, r + r_{c})$$

$$t = 0 \qquad \qquad t = t_{0} < t_{0}$$

$$(\rho, \hat{V}) = (\rho^{0}, V^{0}) \qquad \qquad (\rho, \hat{V}) = (\rho, V)(t_{0})$$
close in natural norms

$$\mathcal{E}^{0}, W^{0} = (\rho^{0}, V^{0}, \mathcal{R}^{0}, r^{0}) \Rightarrow W = (\rho, V, \mathcal{R}, r + r_{c})$$
Say $\rho > 0$, $R = \mathcal{R} + r\mathbf{I}_{n} > 0$

$$\mathcal{E}^{0}, W^{0} \qquad \qquad \mathcal{E}^{0}, W = (\rho, V, \mathcal{R}, r + r_{c})$$

$$t = 0 \qquad \qquad t = t_{0} + t_{0} +$$

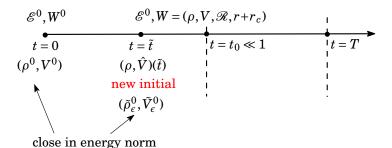
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Say $\rho > 0$, $R = \mathcal{R} + r\mathbf{I}_{n} > 0$

$$\mathcal{E}^{0}, W^{0} \qquad \mathcal{E}^{0}, W = (\rho, V, \mathcal{R}, r + r_{c})$$

$$t = 0 \qquad t = \tilde{t} \qquad |t = t_{0} \ll 1$$

$$(\rho^{0}, V^{0}) \qquad (\rho, \hat{V})(\tilde{t})$$
new initial
$$(\tilde{\rho}_{\varepsilon}^{0}, \tilde{V}_{\varepsilon}^{0})$$
close in energy norm
$$\mathcal{E}^{0} = \int E(\tilde{\rho}_{\varepsilon}^{0}, \tilde{V}_{\varepsilon}^{0})$$

$$\begin{split} \mathcal{E}^0, W^0 &= (\rho^0, V^0, \mathcal{R}^0, r^0) \Rightarrow W = (\rho, V, \mathcal{R}, r + r_c) \\ \text{Say} \quad \rho &> 0, \quad R = \mathcal{R} + r \mathbf{I}_n > 0 \end{split}$$



$$\mathcal{E}^0 = \int E(\tilde{\rho}_{\varepsilon}^0, \tilde{V}_{\varepsilon}^0)$$

Note: ∞ many choices for $\tilde{t} \Rightarrow \infty$ many initial data $(\tilde{\rho}_{\epsilon}^0, \tilde{V}_{\epsilon}^0)$

$$\mathcal{E}^{0}, W^{0} = (\rho^{0}, V^{0}, \mathcal{R}^{0}, r^{0}) \Rightarrow W = (\rho, V, \mathcal{R}, r + r_{c})$$
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$$\mathcal{E}^{0}, W^{0} \qquad \qquad \mathcal{E}^{0}, W = (\rho, V, \mathcal{R}, r + r_{c})$$

$$t = 0 \qquad \qquad t = t_{0} \ll 1$$

$$C.I. \Rightarrow (\rho, \hat{V}) \qquad \qquad C.I. \Rightarrow (\rho, \overline{V})$$

same at t_0

$$\mathcal{E}^{0}, W^{0} = (\rho^{0}, V^{0}, \mathcal{R}^{0}, r^{0}) \Rightarrow W = (\rho, V, \mathcal{R}, r + r_{c})$$
Say $\rho > 0$, $R = \mathcal{R} + r\mathbf{I}_{n} > 0$

$$\mathcal{E}^{0}, W^{0} \qquad \mathcal{E}^{0}, W = (\rho, V, \mathcal{R}, r + r_{c})$$

$$t = \overline{t} \qquad | t = t_{0} \ll 1 \qquad | t = T$$

$$(\tilde{\rho}_{\varepsilon}^{0}, \tilde{V}_{\varepsilon}^{0}), \mathcal{E}^{0} \qquad | (\rho, \overline{V})(t + \tilde{t}) \qquad | (\rho, \overline{V})(t + \tilde{t})$$

$$C.I. \Rightarrow (\rho, \hat{V}) \longrightarrow C.I. \Rightarrow (\rho, \overline{V})$$

